

# EVALUATING ENERGY FLUXES DURING COMPASS DISRUPTIONS

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## KEY RESULTS

- A method to estimate global energy losses during disruptions, based exclusively on magnetic measurements, is proposed
- The method does not require equilibrium reconstruction procedures, making it compatible to eventual real time or monitoring applications
- For a set of COMPASS disruptions the method is compared with data from IR and AXUV pinhole cameras, showing good agreement in terms of overall losses
- The greater magnetic energy available for conversion to heat is verified experimentally and compared to simplified theoretical models
- Preliminary validation against quasi-static MHD simulations is illustrated

## COMPASS

### Key parameters [1]

- Small machine ( $R_w = 0.56$  m,  $\langle b_w \rangle = 0.3$  m) with ITER-like shape (@ IPP CAS 2009-2021)
- Typical toroidal plasma current 160 kA
- Typical TF coil magnetic field at the plasma axis 0.9 – 2.1 T.

### Diagnostics used in this work

- Poloidal, radial and toroidal Mirnov coils, 24 diagnostics per group, equally spaced along the poloidal angle;
- 8 toroidal full flux loops for the measurement of the poloidal flux;
- Diamagnetic flux loop for the toroidal flux measurement (+TF coil current meas.)

## METHODOLOGY

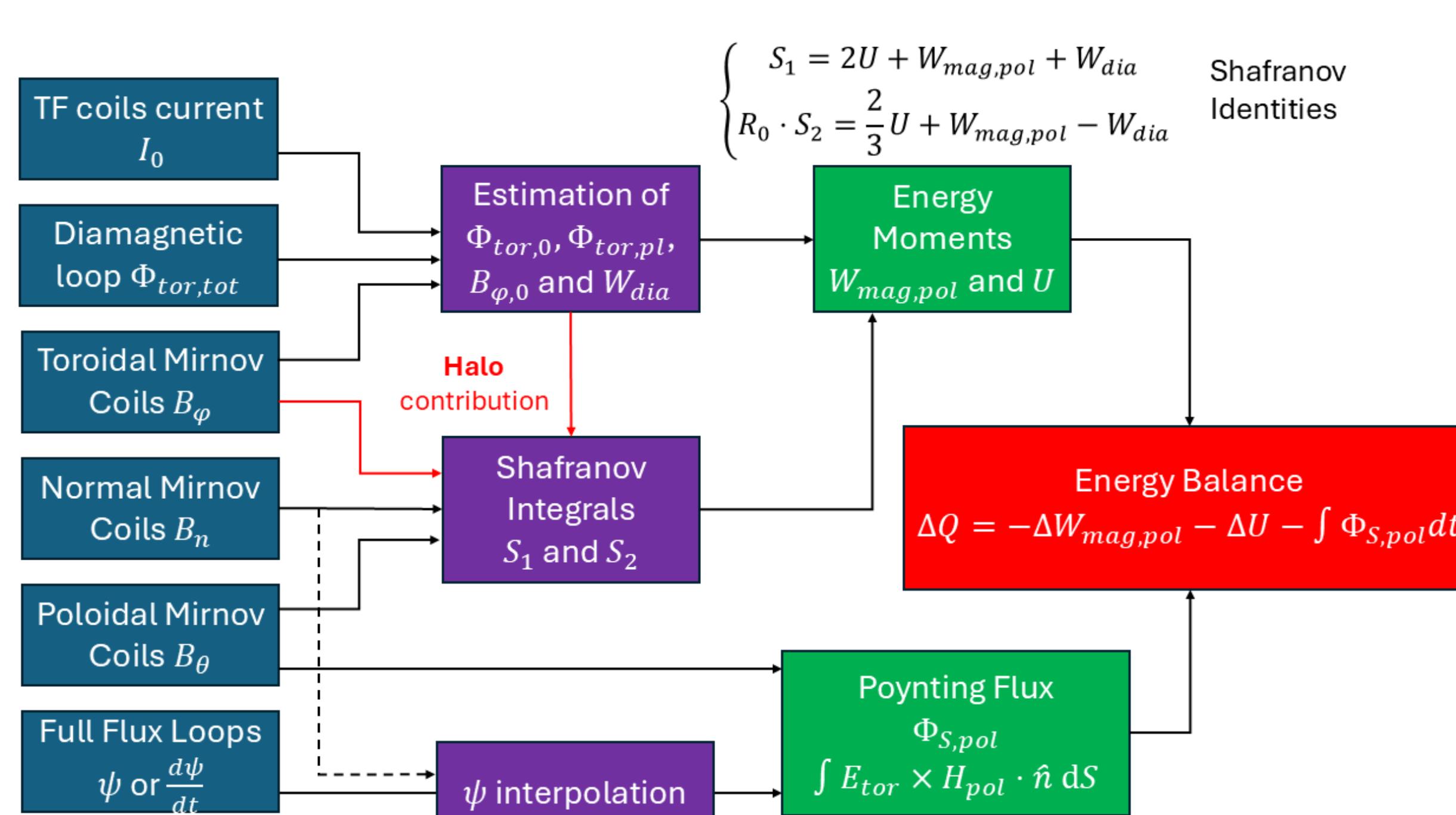


Figure 1: Schematic of the proposed magnetic method for losses estimation.

- Definition of a control volume enclosed by a well-diagnosed Diagnostic Surface;
- Proper discretization of Poynting and Shafranov integrals according to available diagnostics;
- Smart use of normal magnetic field to the Diagnostic Surface to estimate poloidal flux at other locations, choosing automatically non-saturate measurements.

## BENCHMARK WITH OTHER DIAGNOSTICS

- Benchmark against sum of conducted and radiated heat as measure by IR camera and AXUV pinhole diodes respectively;
- IR camera points to the divertor region → only downward VDEs considered for benchmark;
- Toroidal Mirnov coils signals need lots of processing, dedicated experiments with tuned range or toroidal Rogowski coils may help.

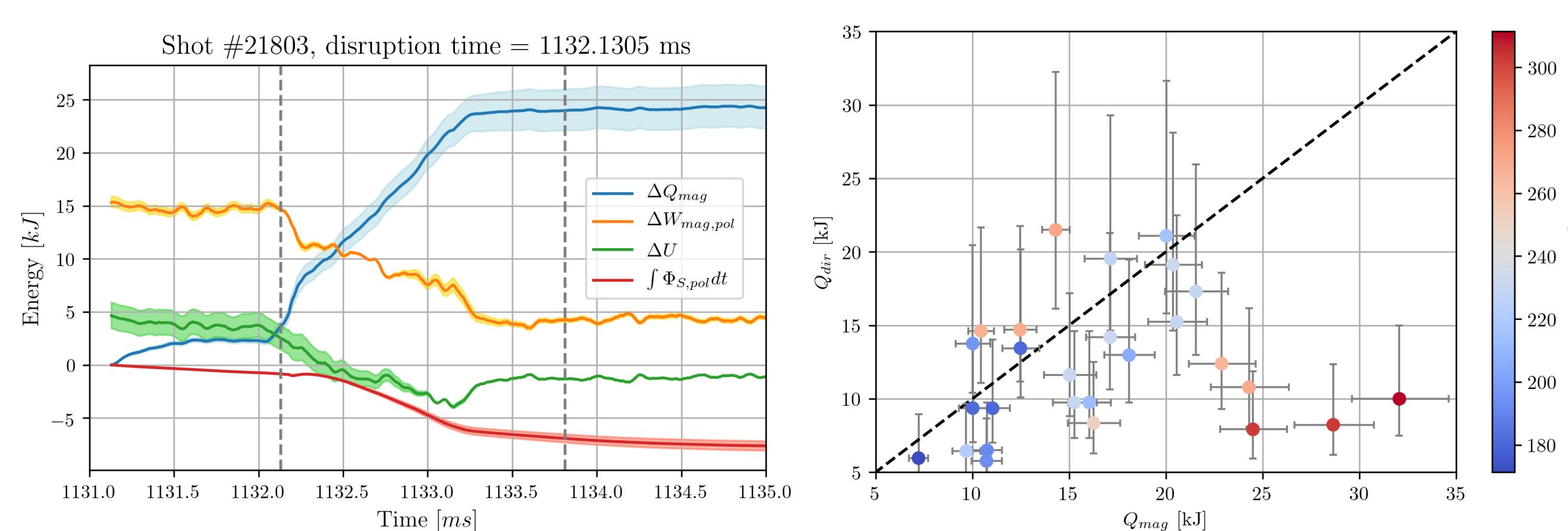


Figure 2: (a) Example of time integrated Poynting integral and energy moment variation estimated magnetically during a COMPASS disruption; (b) Parity plot comparing proposed magnetic method and direct conducted and radiated heat measurements.

## BENCHMARK WITH THEORETICAL PREDICTIONS

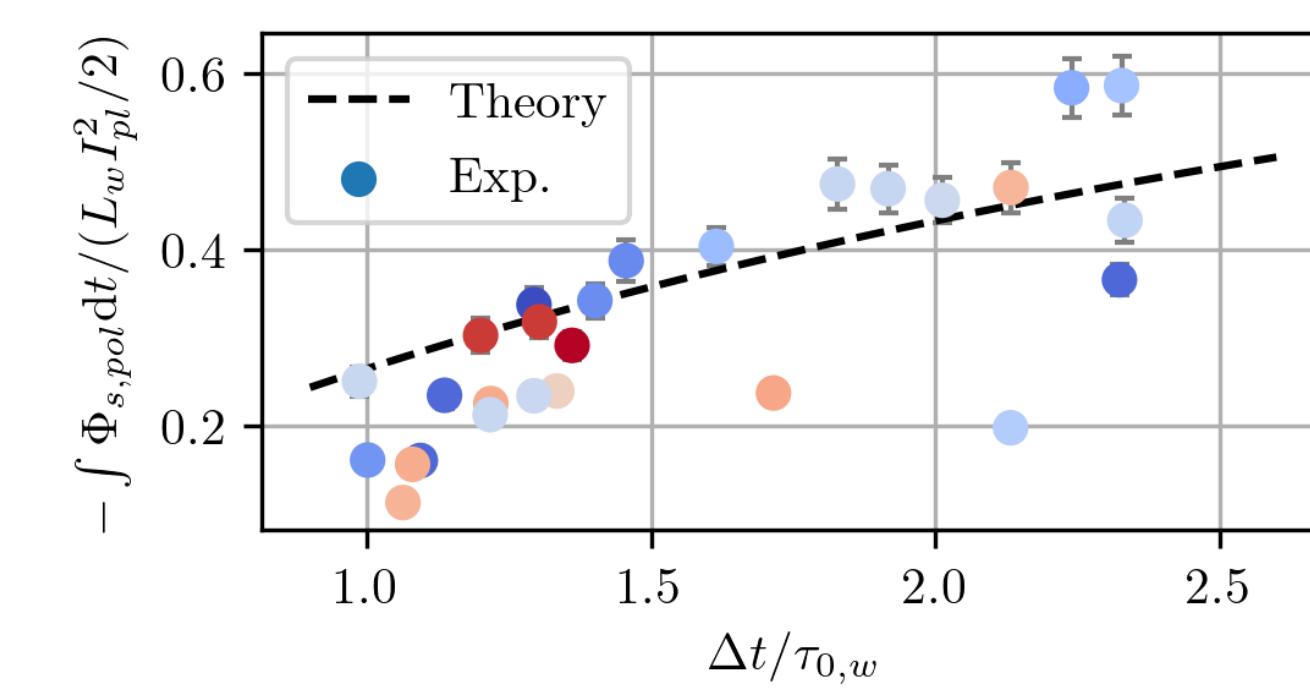


Figure 3: Magnetic energy influx as function of disruption duration.

- Longer CQ duration → larger the fraction of magnetic energy penetrating inside the vessel → more magnetic energy is converted to heat within the plasma [2]
- $L_w = 49.6 \mu\text{H}$ ,  $\tau_{0,w} = 0.84 \pm 0.21 \text{ ms}$
- Analytical expression for the normalized energy influx:

$$\frac{2}{(\Delta t / \tau_{0,w})^2} \left[ 1 - \frac{\Delta t}{\tau_{0,w}} + \frac{1}{2} \left( \frac{\Delta t}{\tau_{0,w}} \right)^2 - \exp \left( -\frac{\Delta t}{\tau_{0,w}} \right) \right].$$

## BENCHMARK WITH SIMULATIONS

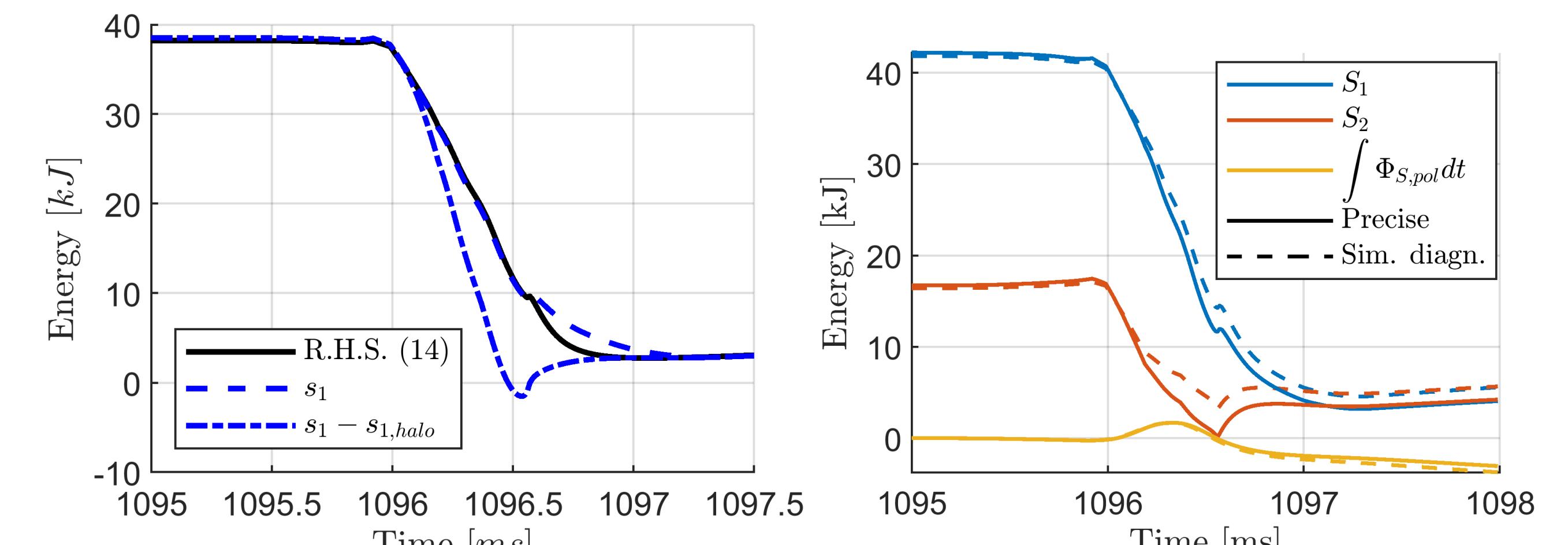


Figure 4: (a) Relative weight of halo contribution in the Shafranov identity for  $S_1$ , (b) Error introduced considering a discretization of the surface integrals at Mirnov coil locations.

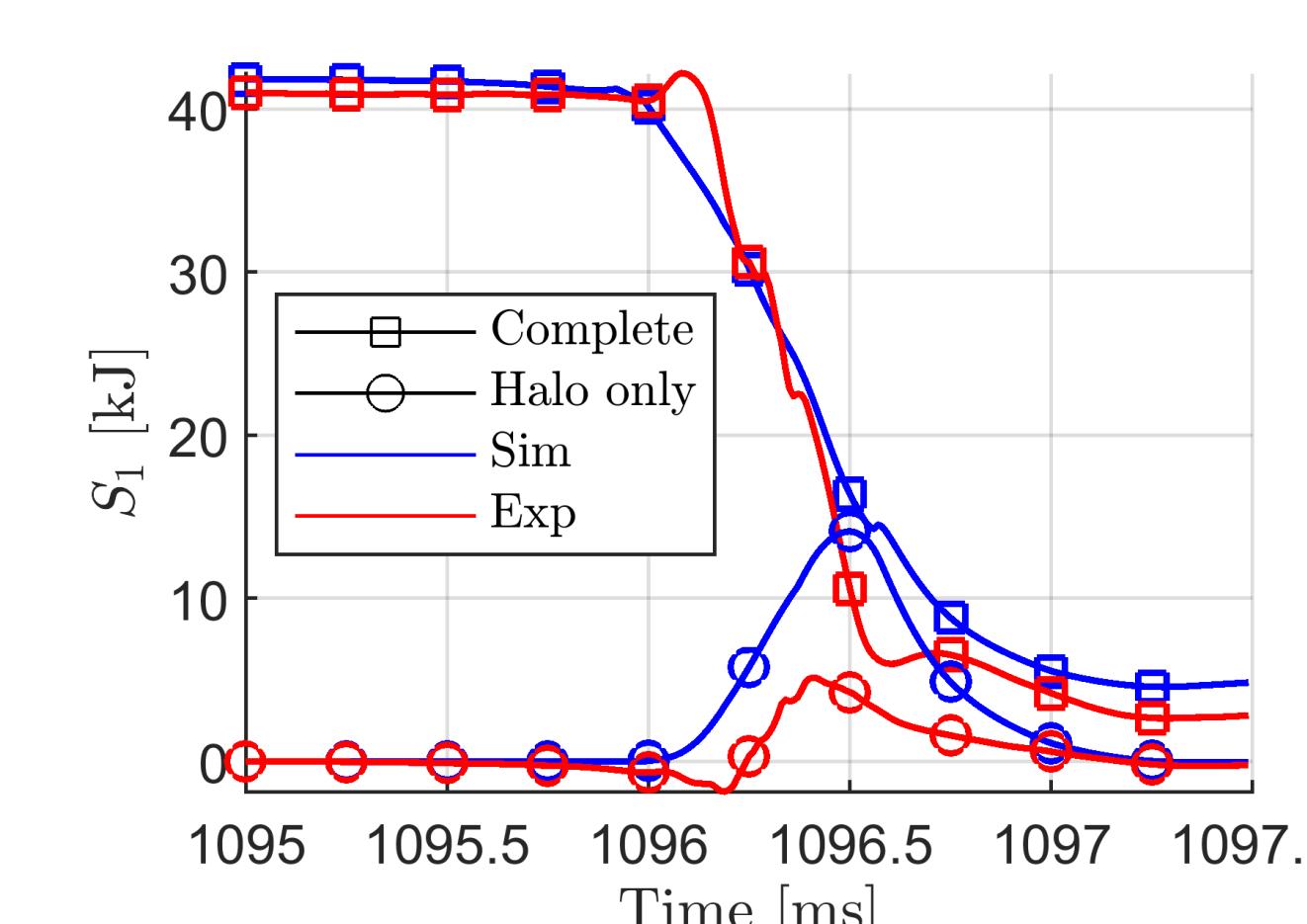


Figure 5: Comparison of Shafranov integrals in experiment #19172 and respective CarMa0NL simulation. Halo contribution is highlighted.

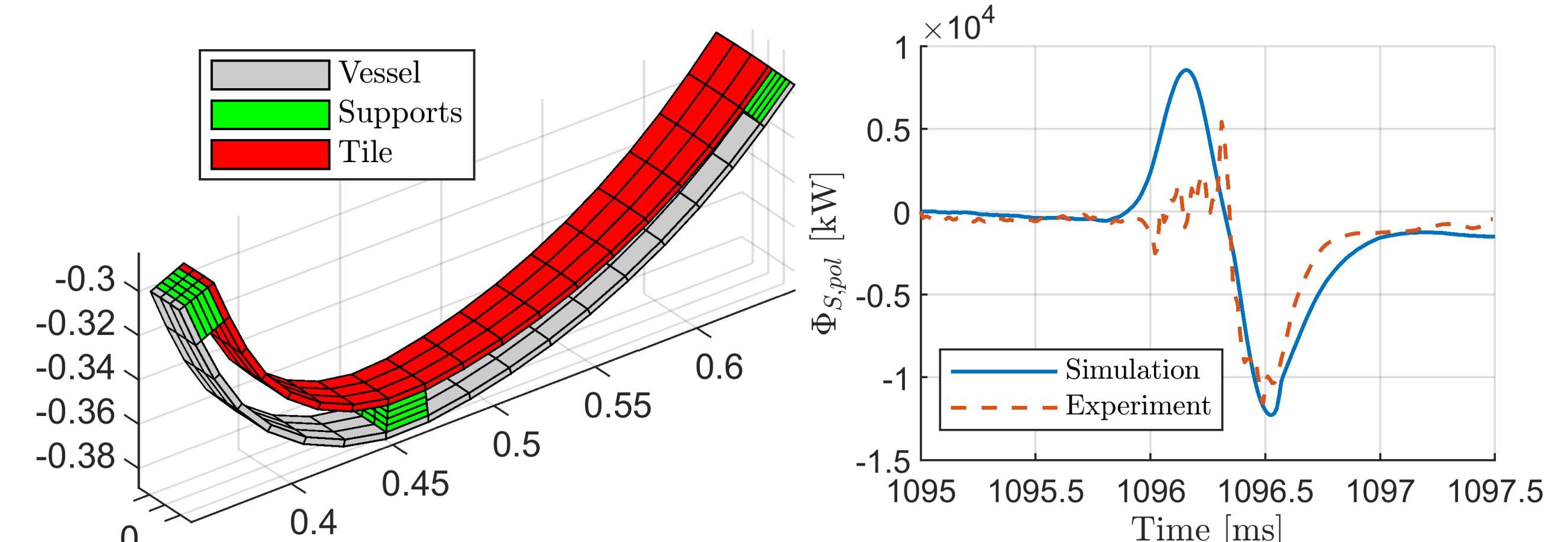


Figure 6: Approximate tiles representation to study halo current patterns.

Figure 7: Flux of Poynting vector across Diagnostic Surface.

- Mass-less MHD models share the fundamental assumption of the proposed method → CarMa0NL simulations [5];
- Halo currents are important in the correct determination of poloidal magnetic energy and plasma thermal energy;
- COMPASS diagnostics set up has sufficient resolution for determining correctly the Shafranov and Poynting integrals;

## REFERENCES

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